



## Table of Constants

*(All constants are in SI)*

Parameter	Symbol	Value
Gravitational constant	$G$	$6.67 \times 10^{-11} \text{ N m}^2 \text{ kg}^{-2}$
Plank constant	$h$	$6.63 \times 10^{-34} \text{ J s}$
	$\hbar$	$1.05 \times 10^{-34} \text{ J s}$
Speed of light	$c$	$3.00 \times 10^8 \text{ m s}^{-1}$
Solar Mass	$M_{\odot}$	$1.99 \times 10^{30} \text{ kg}$
Solar radius	$R_{\odot}$	$6.96 \times 10^8 \text{ m}$
Solar luminosity	$L_{\odot}$	$3.83 \times 10^{26} \text{ W}$
Apparent solar magnitude ( $V$ )	$m_{\odot}$	-26.8
Solar constant	$b_{\odot}$	$1.37 \times 10^3 \text{ W m}^{-2}$
Mass of the Earth	$M_{\oplus}$	$5.98 \times 10^{24} \text{ kg}$
Radius of the Earth	$R_{\oplus}$	$6.38 \times 10^6 \text{ m}$
Mean density of the Earth	$\rho_{\oplus}$	$5 \times 10^3 \text{ kg m}^{-3}$
Gravitational acceleration at sea level	$g$	$9.81 \text{ m s}^{-2}$
Tropical year		365.24 days
Sidereal year		365.26 days
Sidereal day		86164 s
Inclination of the equator with respect to the ecliptic	$\epsilon$	23.5°
Parsec	$pc$	$3.09 \times 10^{16} \text{ m}$
Light year	$ly$	$9.46 \times 10^{15} \text{ m}$
Astronomical Unit	$AU$	$1.50 \times 10^{11} \text{ m}$
Solar distance from the center of the Galaxy	$R$	$8 \times 10^3 pc$
Hubble constant	$H$	$75 \text{ kms}^{-1} \text{ M pc}^{-1}$
Mass of electron	$m_e$	$9.11 \times 10^{-31} \text{ kg}$
Mass of proton	$m_p$	$1.67 \times 10^{-27} \text{ kg}$
Central wavelength of V-band	$\lambda$	550 nm
Refraction of star light at horizon		34'
	$\pi$	3.1416

*Useful mathematical formula:*  
 $\ln(1 + x) \sim x \quad \text{for } x \rightarrow 0$

**Problem 16: High Altitude Projectile (45 points)**

A projectile which starts from the surface of the Earth at the sea level is launched with the initial speed of  $v_0 = \sqrt{GM/R}$  and with the projecting angle (with respect to the local horizon) of  $\theta = \frac{\pi}{6}$ .  $M$  and  $R$  are the mass and radius of the Earth respectively. Ignore the air resistance and rotation of the Earth.

- Show that the orbit of the projectile is an ellipse with a semi-major axis of  $a = R$ .
- Calculate the highest altitude of the projectile with respect to the Earth surface (in the unit of the Earth radius).
- What is the range of the projectile (distance between launching point and falling point) in the units of the earth radii?
- What is eccentricity ( $e$ ) of this elliptical orbit?
- Find the time of flight for the projectile.

**Problem 17: Apparent number density of stars in the Galaxy (45 points)**

Let us model the number density of stars in the disk of Milky Way Galaxy with a simple exponential function of  $n = n_0 \exp\left(-\frac{r-R_0}{R_d}\right)$ , where  $r$  represents the distance from the center of the Galaxy,  $R_0$  is the distance of the Sun from the center of the Galaxy,  $R_d$  is the typical size of disk and  $n_0$  is the stellar density of disk at the position of the Sun. An astronomer observes the center of the Galaxy within a small field of view. We take a particular type of Red giant stars (red clump) as the standard candles for the observation with approximately constant absolute magnitude of  $M = -0.2$ ,

- Considering a limiting magnitude of  $m = 18$  for a telescope, calculate the maximum distance that telescope can detect the red clump stars. For simplicity we ignore the presence of interstellar medium so there is no extinction.
- Assume an extinction of  $0.7 \text{ mag/kpc}$  for the interstellar medium. Repeat the calculation as done in the part (a) and obtain a rough number for the maximum distance these red giant stars can be observed.
- Give an expression for the number of these red giant stars per magnitude within a solid angle of  $\Omega$  that we can observe with apparent magnitude in the range of  $m$  and  $m + \Delta m$ , (i.e.  $\frac{\Delta N}{\Delta m}$ ). Red giant stars contribute fraction  $f$  of overall stars. In this part assume no extinction in the interstellar medium as part (a). Assume the size of the disk is larger than maximum observable distance.