

VIII Международная астрономическая олимпиада
VIII International Astronomy Olympiad
VIII:e Internationella Astronomiolympiaden

Швеция, Стокгольм

2 – 8. 10. 2003

Stockholm, Sweden

Тезисы решений задач практического тура

Practical round. Sketches for solutions

язык	<i>English</i>
language	

Groups A,B

7. Measuring the speed of light

7.1. Let A be Saturn's orbital radius, B be Saltis's orbital radius and c be the velocity of light. It takes light the time A/c to reach the Sun from Saturn. At opposition it takes light $(A-B)/c$ to reach Saltis from Saturn, and at conjunction $(A+B)/c$. This means that the difference in the timings from the tabulated values will be $A/c - (A-B)/c = B/c$ at opposition (i.e. the eclipse will be seen earlier) and $A/c - (A+B)/c = -B/c$ at conjunction (i.e. the eclipse will be observed later than tabulated values).

Looking at the differences in Table 1, one sees that the spread in value is greater than the estimated timing error of 0.03 pinit. Thus the spread is mainly due to Saturn not being exactly at opposition/conjunction at the time of eclipse. Thus the extreme value is the best approximation, 6.23 pinit. Since the three opposition values are ranging from 6.12–6.23 pinit, the true opposition difference is probably not larger than 6.30 pinit. Thus a sensible estimated error is ~ 0.1 pinit. We ignore the conjunction values since they are further from true conjunction, probably due to difficulties in observing Saturn near the Sun.

With a time difference $B/c = 6.23 \pm 0.1$ pinit, and B defined to 10^9 seter, we readily compute $c = (1.61 \pm 0.03) \times 10^8$ seter/pinit.

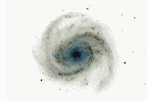
Table 1 Eclipses of Titan by Saturn

Tabulated (pinit)	Celesta (pinit)	Difference (pinit)
456.47	450.32	6.15
18.50	12.28	6.23
821.41	815.29	6.12
444.70	450.85	-6.14
615.43	621.52	-6.08
791.94	798.02	-6.08

7.2. The difference between the largest and smallest distance between Saltis and Earth is equal to 2 Earth orbital radii. From Fig. 1 we estimate the time difference to 5.3 pinit. With c estimated as above, we get $5.3c = 8.54 \cdot 10^8$ seter. Thus $1 \text{ AU} = 4.3 \cdot 10^8$ seter.

7.3. $1 \text{ AU} = 149.6 \cdot 10^6 \text{ km} = 4.3 \cdot 10^8$ seter gives $1 \text{ seter} = 348 \text{ meter}$. With $c = 2.998 \cdot 10^8 \text{ m/s} = 1.61 \cdot 10^8$ seter/pinit we have $1 \text{ pinit} = 187 \text{ seconds}$.

7.4. From Fig. 1 we estimate the interval between oppositions to $2.3 \cdot 10^5$ pinit = 1.4 yr. Since the period of Earth is 1 yr, we use the formula for the synodic period ($1/T = 1/P_1 - 1/P_2$) to get the period of 3.7 yr. Alternatively, one may use the orbital radius of Saltis $a = 348 \cdot 10^9 \text{ m} = 2.33 \text{ AU}$ and Kepler's law, $p^2 = a^3$, to get $p = a^{3/2} = 2.33^{3/2} \text{ yr} = 3.6 \text{ yr}$.



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Group B

8. Estimating the mass of Saturn

8.1. If Saturn's ring was a rigidly rotating body, its outer edge would be faster than its inner edge, and the spectrum would look like Fig. 3.

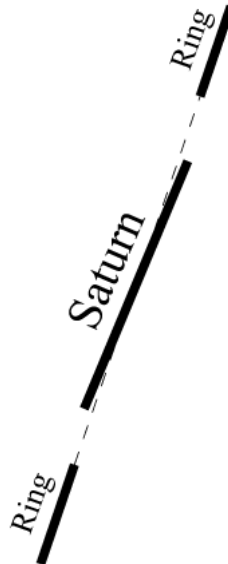


Figure 3 The rotation curve for Saturn with a rigid ring. The rotation curve of the ring is a straight line, with the outer edge rotating faster than the inner, and with the whole ring approximately at the Kepler velocity.

8.2. We start by determining the scale of the spectrum. The distance between the two Na I lines in the spectrum of Fig. 2 is 84 mm. Since we know the two D lines of Na I to differ by 0.59 nm, we have the spectral scale $0.59/84 \text{ nm/mm} = 0.007024 \text{ nm/mm}$.

Light from the solar spectrum is received by the reflecting surface of Saturn at a Doppler shift depending on the velocity of the surface with respect to the Sun. With respect to an observer, the light reflected on the surface is then again Doppler shifted with the velocity between the observer and the surface. From the small shadow of Saturn on its ring in Fig. 1, we note that the angle between the Earth and the Sun, as seen from Saturn, is small. Thus the Doppler shift differences between different parts of the surface (and the rings) are approximately just doubled, i.e. $\Delta\lambda/\lambda_0 = 2v/c$.

However, we observe Saturn at an inclination, meaning that we observe only the component of the velocity pointing in our direction. Since the ring system is circular in shape and planar, we can measure the projection factor by measuring the dimensions of the projected ring. The major axis of the ring

measured in Fig. 1 is 116 mm, while the minor axis is 51 mm. By Pythagoras's theorem (see Fig. 4), this gives the projection factor

$$(116^2 - 51^2)^{1/2} / 116 = 0.90.$$

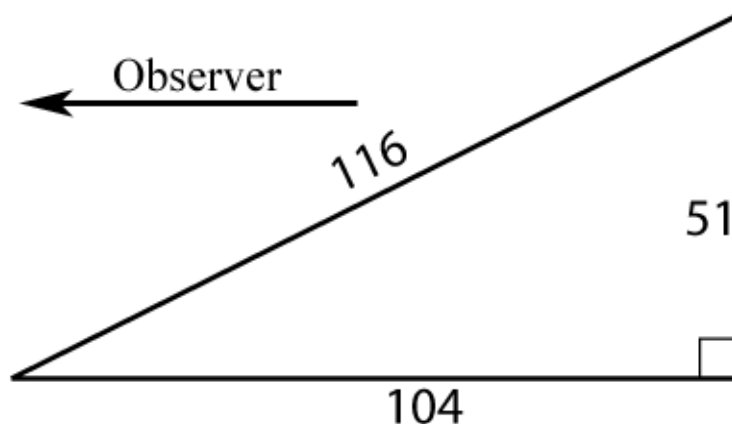


Figure 4 If the velocity is 116, then the projected velocity component in the direction of the observer is 104, that is, the projection factor is $104/116 = 0.90$.

For the spectrum of Fig. 2, this translates to the velocity scale

$$0.5 \times 0.007024 \times c / (589.3 \times 0.90) \text{ mm}^{-1} = 1.98 \text{ km/s mm}^{-1}.$$

We can now calculate the difference in velocity between the different edges of the planetary disc. The difference is about 10 mm, meaning 20 km/s. During the period $10.66 \text{ h} = 38000 \text{ s}$, the surface thus travels $10 \times 38000 \text{ km} = 380000 \text{ km}$, which is the equatorial circumference of Saturn. Dividing by π we get the diameter 121000 km.

8.3. Since the equatorial diameter of the planet is 121000 km but 18 mm in the spectrum of Fig. 2, the image scale is $121000/18 \text{ km/mm} = 6700 \text{ km/mm}$. Looking at the inner part of the ring, the displacement of an absorption line between the two sides is 20 mm, corresponding to the velocity difference $20 \times 1.98 \text{ km/s} = 39.7 \text{ km/s}$. The diameter of the inner ring is 26 mm in Fig. 2, meaning $26 \times 6700 \text{ km} = 175000 \text{ km}$. Using Kepler's third law,

$$\left(\frac{P}{\text{yr}} \right)^2 = \frac{(a/\text{AU})^3}{M/M_{\text{Sun}}},$$

where $\text{yr} = 3.16 \times 10^7 \text{ s}$, $\text{AU} = 1.496 \times 10^8 \text{ km}$ and $M_{\text{Sun}} = 1.99 \times 10^{30} \text{ kg}$. The circumference of the ring of diameter 175000 km is 549000 km, giving the period $549000 / (39.7 \times 0.5) = 27700 \text{ s}$. Thus Saturn's mass is calculated as

$$M = \frac{(a/\text{AU})^3}{(P/\text{yr})^2} M_{\text{Sun}} = \frac{(8.74 \times 10^4 / 1.496 \times 10^8)^3}{(2.77 \times 10^4 / 3.16 \times 10^7)^2} 1.99 \times 10^{30} \text{ kg} = 5.2 \times 10^{26} \text{ kg}.$$

What is the expected accuracy of this number? Assuming our measurements on paper are accurate to $\pm 0.5 \text{ mm}$, one can derive the formal error $5 \times 10^{25} \text{ kg}$. The mass of Saturn, as measured accurately by Voyager, is $5.685 \times 10^{26} \text{ kg}$, consistent with our value.